Solution of self-consistent equations for the N³LO nuclear energy density functional in spherical symmetry. The program HOSPHE (v1.02) 

B.G. Carlsson a, J. Dobaczewski a,b,*, J. Toivanen a, P. Veselý a

a Department of Physics, P.O. Box 35 (YFL), FI-40014 University of Jyväskylä, Finland
b Institute of Theoretical Physics, University of Warsaw, Hoża 69, PL-00681 Warsaw, Poland

Abstract

We present solution of self-consistent equations for the N³LO nuclear energy density functional. We derive general expressions for the mean fields expressed as differential operators depending on densities and for the densities expressed in terms of derivatives of wave functions. These expressions are then specified to the case of spherical symmetry. We also present the computer program HOSPHE (v1.02), which solves the self-consistent equations by using the expansion of single-particle wave functions on the spherical harmonic oscillator basis.

Program summary

Program title: HOSPHE (v1.02)
Catalogue identifier: AEGK_v1_0
Program summary URL: http://cpc.cs.qub.ac.uk/summaries/AEGK_v1_0.html
Program obtainable from: CPC Program Library, Queen’s University, Belfast, N. Ireland
No. of lines in distributed program, including test data, etc.: 45 809
No. of bytes in distributed program, including test data, etc.: 290 514
Distribution format: tar.gz
Programming language: Fortran-90
Computer: PCs and workstations
Operating system: Linux
RAM: 50 MB
Classification: 17.22
External routines: LAPACK (http://www.netlib.org/lapack/), BLAS (http://www.netlib.org/blas/)
Nature of problem: The nuclear mean-field methods constitute principal tools of a description of nuclear states in heavy nuclei. Within the Local Density Approximation with gradient corrections up to N³LO [1], the nuclear mean-field is local and contains derivative operators up to sixth order. The locality allows for an effective and fast solution of the self-consistent equations.
Solution method: The program uses the spherical harmonic oscillator basis to expand single-particle wave functions of neutrons and protons for the nuclear state being described by the N³LO nuclear energy density functional [1]. The expansion coefficients are determined by the iterative diagonalization of the mean-field Hamiltonian, which depends non-linearly on the local neutron and proton densities.
Restrictions: Solutions are limited to spherical symmetry. The expansion on the harmonic-oscillator basis does not allow for a precise description of asymptotic properties of wave functions.
Running time: 50 seconds of CPU time for the ground-state of 208Pb described by using N₀ = 50 as the maximum harmonic-oscillator shell included in the basis.
References:

© 2010 Elsevier B.V. All rights reserved.
1. Introduction

The nuclear mean-field methods constitute principal tools of a description of nuclear states in heavy nuclei [1]. Their applicability to nuclei is interpreted within the Hohenberg–Kohn [2] and Kohn–Sham [3] theorems involving the nuclear energy density functional (EDF). Recently, we formulated the NLO nuclear EDF with gradient corrections up to sixth order [4]. The present study presents practical formulation of the method, allowing for a solution of the corresponding self-consistent equations. We also present the computer program HOSPE (v1.02), which solves the self-consistent equations by using the expansion of single-particle wave functions on the spherical harmonic oscillator (HO) basis.

The paper is organized as follows. In Section 2, we present concise review of the method. In Section 3, we give general forms of the NLO potentials, fields, and densities, which are then in Section 4 specified to the case of spherical symmetry. Sections 5–10 describe the structure, installation, and test runs of the code HOSPE (v1.02), and Section 11 concludes our study.

2. Overview of the method

This introductory section is intended as a guide to the subsequent sections, where more detailed derivations and results are presented. Here, we use abbreviated notations so as to give a brief outline of the method, while referring the reader to the following sections for details.

The mean-field eigenvalue equation is obtained by considering the variation of the energy with the condition of the single-particle wave functions $\phi_i(r, \sigma)$ to be normalized to unity,

$$\frac{\delta}{\delta \phi_i^*(r', \sigma')}(E - \sum_i \epsilon_i \int |\phi_i|^2 d^3r) = 0.$$

The potential energy is expressed as the EDF of Ref. [4], that is,

$$E = \int \sum_{\alpha \beta} C_{\alpha \beta}^0 T_{\alpha \beta}^0(r) d^3r,$$

where the grouped indices, such as the Greek indices $\alpha = [n_{\alpha}, l_{\alpha}, \sigma_{\alpha}, J_{\alpha}]$ or $\beta = [n_{\beta}, l_{\beta}, \sigma_{\beta}, J_{\beta}]$ and the Roman indices $\alpha = [m_{\alpha}]$, denote all the quantum numbers of the local primary $\rho_\alpha(r)$ and secondary $\rho_{\alpha,\beta}(r)$ densities [4]. In Eq. (2), $T_{\alpha \beta}^0(r)$ denotes terms of the functional,

$$T_{\alpha \beta}^0 = \rho_\beta \rho_{\alpha} \right\} 0 = \left[ \rho_\beta \left[ D_{\alpha} \rho_\alpha \right] J_{\alpha} \right]_0,$$

where $C_{\alpha \beta}$ denote the coupling constants, $D_{\alpha}$ denote the higher order derivative operators [4], and the sum runs over all terms of the functional. Although not shown explicitly, the sum contains both isoscalar and isovector terms. At present we have neglected neutron–proton mixing, which means that only the $m_t = 0$ component of the isovector densities are present. The convention adopted is such that the isovector contributions to the densities are taken as neutron minus proton densities and isoscalar densities are sums of neutron and proton densities.

For each term of the functional, the variation with respect to the local densities, followed by the integration by parts and recoupling, gives

$$\delta \int T_{\alpha \beta}^0 d^3r = \int \left[ \delta \rho_\beta \left[ D_{\alpha} \rho_\alpha \right] J_{\alpha} \right]_0 + (-1)^{J_{\alpha} - J_{\beta}} \left[ \delta \rho_\alpha \left[ D_{\beta} \rho_\beta \right] J_{\beta} \right]_0 d^3r.$$

The higher order derivative operators $K_{n\ell}$ [4] are built by coupling the relative-momentum operators $k = \frac{1}{2} (V^\beta - V^\alpha)$, where we have used the subscripts to indicate on which function the operators act upon. This allows us to perform the variation with respect to $\phi_i^*$.

$$\frac{\delta}{\delta \phi_i^*} \int T_{\alpha \beta}^0 d^3r = \left[ K_{n\ell}^i \left[ \sigma v_i \phi_i \right] J_{\alpha} \left[ D_{\beta} \rho_\beta \left] J_{\beta} \right]_0 \right. + (-1)^{J_{\alpha} - J_{\beta}} \left. \left[ K_{n\ell}^i \left[ \sigma v_i \phi_i \right] J_{\beta} \left[ D_{\alpha} \rho_\alpha \left] J_{\alpha} \right]_0 \right. \right.$$

After performing the variation, the integral above was partially integrated so that the derivatives would not act on the variation of $\phi_i^*$. Therefore, the $K_{n\ell}$ operators are built by coupling the relative-momentum operators $k' = \frac{1}{2} (V^\beta + V^\alpha)$, where $V$ acts on all the functions of position standing to the right.

The operator on the right-hand side of Eq. (6) is a formal expression for the mean-field operator. All what remains to be done is to disentangle the gradients $V^\alpha$ and $V$ from one another – this procedure is performed in Eqs. (54)-(60) below. Finally, the mean-field operator $h(\rho)$ acquires the form:

$$h(\rho) = \sum_{\alpha \beta} \left[ U_{\alpha} \left[ D_{\beta} \sigma v_\alpha \right] J_{\beta} \right]_0,$$

where the differential operators $D_{n\ell}$ and Pauli matrices $\sigma v_\alpha$ act on the single-particle wave functions, and the potentials $U_{\alpha}(r)$ are linear combinations of the secondary densities:

$$U_{\alpha} = \sum_{\alpha \beta} \chi_{\alpha \beta}^0 \left[ D_{\beta} \sigma v_\alpha \right] J_{\beta}.$$

The coefficients $\chi_{\alpha \beta}^0$ can be derived by using the recoupling rules presented in Section 3.3. An alternative method, which was also used when building the code HOSPE (v1.02), was to construct the fields by starting from Eq. (6) and putting them equal to those of Eq. (7). This gives a linear system of equations that can be solved for the unknown coefficients $\chi_{\alpha \beta}^0$. At NLO, only 1494 such coefficients are needed, so they can easily be precalculated and stored.

It is now clear, that the key operators in the mean field are given by

$$F_{\alpha \beta} = \left[ \left[ \rho_{\alpha \beta} \right] J_{\beta} \left[ D_{n\ell} \sigma v_\alpha \right] J_{\beta} \right]_0,$$

and their matrix elements in the single-particle basis $\phi_i(r, \sigma)$ read,

$$F_{\alpha \beta}^i = \sum_{\alpha' \beta'} \chi_{\alpha \beta}^0 \phi_i \left[ \delta \rho_\beta \left[ D_{\alpha} \rho_\alpha \right] J_{\alpha} \right]_0 d^3r,$$

Then, the mean-field matrix elements can be written as the following sum:

$$\phi_i \left[ h(\rho) \phi_i \right] = \sum_{\alpha \beta; \delta; \gamma} C_{\alpha \beta}^0 \chi_{\alpha \beta}^0 \left[ D_{\delta} \sigma v_\alpha \right] J_{\delta} F_{\alpha \beta}^i.$$
When constructing potentials [8], we need expressions to calculate all secondary densities. These can be written as [see Eqs. (75)–(77)]:
\[
\rho_{d,\delta,JM} = [D_d\rho_{k}]_{JM} = \sum_{bb/W} A^{bb',W}_{d,\delta,J} \rho_{\nu_{b'}}_{JM},
\]
(12)
with
\[
\rho_{\nu_{b'}}_{JM}(r_1) = \left\{ \left[D^{(1)}_b D^{(2)}_b \right]_{W} \rho_{\nu_{b'}}(r_1, r_2) \right\}_{r_2=r_1},
\]
(13)
where the superscripts on the derivative operators indicate on which coordinate they act. The coefficients \(A\) can be obtained by explicitly constructing the left- and right-hand sides of Eq. (12), which gives a linear system of equations in derivatives of the density matrix that can be solved for the unknown coefficients \(A^{bb',W}_{d,\delta,J}\).

At \(N^3LO\), only 3138 such coefficients are needed, so they can easily be precalculated and stored. In Section 3.6 we also show how to derive these coefficients by using the recoupling rules and in Section 4.2 we give the expressions for densities in the spherical HO basis.

3. General forms of the \(N^3LO\) potentials, fields, and densities

3.1. Building blocks

We begin by recalling definitions that are used to construct operators and densities in the spin and position space. The basic building blocks are given as in Eqs. (8)–(9) of [4], i.e.,
\[
\sigma = 0, 0 = \sigma_0,
\]
(14)
\[
\sigma = 1, \mu = -1, 0, 1 = -i\left\{ \frac{1}{2}(\sigma_x - i\sigma_y), \sigma_z, \frac{1}{2}(\sigma_x + i\sigma_y) \right\},
\]
(15)
\[
\nabla = -i\left\{ \frac{1}{2}(\nabla_x - i\nabla_y), \nabla_z, \frac{1}{2}(\nabla_x + i\nabla_y) \right\},
\]
(16)
\[
k = \frac{1}{2i}(\nabla - \nabla^\dagger),
\]
(18)
where \(k\) is the relative momentum operator.

All possible \(N^3LO\) differential operators \(D_{nLM}\), which can be built of gradients (16), are given in Table 1 of Ref. [4], where \(n\) is the order of the operator and \(L\) is its rank with magnetic projection \(M\). Exactly in the same way, in Ref. [4] we defined the operators \(K_{nLM}\), which are spherical tensors built of the relative momentum operators \(k\) (17).

Hermitian-conjugation properties of the building blocks read:
\[
\sigma_{\nu\mu}^\dagger = Q_\sigma (-1)^{1-\mu} \sigma_{\nu,-\mu} \quad \text{for} \quad Q_\sigma = +1,
\]
(19)
\[
\nabla_{\nu\mu}^\dagger = Q_\nu (-1)^{1-\mu} \nabla_{\nu,-\mu} \quad \text{for} \quad Q_\nu = -1,
\]
(20)
\[
k_{\mu\mu}^\dagger = Q_k (-1)^{1-\mu} k_{\mu,-\mu} \quad \text{for} \quad Q_k = +1.
\]
(21)
For any pair of commuting operators \(A_{\lambda\mu}\) and \(B_{\lambda\mu}\) that have the following hermitian-conjugation properties:
\[
A_{\lambda\mu}^\dagger = Q_A (-1)^{1-\mu} A_{\lambda,-\mu},
\]
(22)
\[
B_{\lambda\mu}^\dagger = Q_B (-1)^{1-\mu} B_{\lambda,-\mu},
\]
(23)
the operator \(C_{LM}\) built by the angular momentum coupling,
\[
C_{LM} = [A_{\lambda},B_{\lambda'}]\sub{LM} = \sum_{\mu\mu'} C_{LM,\mu\mu'} A_{\lambda\mu} B_{\lambda'\mu'},
\]
(24)
behaves under the hermitian conjugation as:
\[
C_{LM}^\dagger = Q_C (-1)^{1-M} C_{L,-M} \quad \text{for} \quad Q_C = Q_A Q_B.
\]
(25)
As a consequence, we have
\[
D_{nLM}^+ = (-1)^n (-1)^{1-M} D_{nLM},
\]
(26)
\[
K_{nLM}^+ = (-1)^{L-M} K_{nLM},
\]
(27)
We note that the gradient operators (16) and (17) obey the Biedenharn–Rose phase conventions of
\[
\nabla_{\nu\mu}^+ = (-1)^{1-\mu} \nabla_{\nu,1-\mu},
\]
(28)
\[
K_{1\mu}^+ = (-1)^{1-\mu} k_{1,-\mu},
\]
(29)
which gives
\[
D_{nLM}^+ = (-1)^{1-M} D_{nLM},
\]
(30)
\[
D_{nLM}^+ = (-1)^{n} D_{nLM},
\]
(31)
\[
D_{nLM}^T = (-1)^{n} D_{nLM},
\]
(32)
and
\[
K_{nLM}^+ = (-1)^{n} (-1)^{1-M} K_{nLM},
\]
(33)
\[
K_{nLM}^+ = (-1)^{n} K_{nLM},
\]
(34)
\[
K_{nLM}^T = (-1)^{n} K_{nLM},
\]
(35)
where superscript \(T\) denotes the transposed operator.

3.2. Potentials

The potential energy to be varied over the wave functions is given in Ref. [4] and reads
\[
\mathcal{E} = \int d^3 r \mathcal{H}(r),
\]
(36)
for
\[
\mathcal{H}(r) = \sum_{n'l'v'l'MnLMJ} C_{n'l'v'J}^{nlvJ}[\rho_{n'l'v'l'M}(r) [D_{n'M} \rho_{nLMJ}(r)]_{0}]
\]
\[
\times \left[ D_{n'M} \rho_{nLMJ}(r) \right]_{J'=1-M},
\]
(37)
where \(C_{n'l'v'J}^{nlvJ}\) are the coupling constants and \(\rho_{nLMJ}(r)\) are the primary densities:
\[
\rho_{nLMJ}(r) = [K_{nl}\rho_{\nu}(r \cdot r')]_{J'=1-M},
\]
(38)
which are built by acting with the relative momentum operators \(K_{nLM}\) on the scalar \((\nu = 0)\) and vector \((\nu = 1)\) non-local densities:
\[
\rho_{\nu}(r \cdot r') = \sum_{\sigma'\sigma} \rho(\sigma \nu \sigma' \nu')(\sigma') |\sigma_{\nu \mu}| \sigma.
\]
(39)
Note that the sum in Eq. (37) runs over the indices ordered in a specific way, defined in Ref. [4], namely,
\[
\{n'l'v'J'\} \leq \{nLMJ\}.
\]
(40)
We first vary \(\mathcal{E}\) over the densities and then, in Section 3.3, we vary densities over the wave functions, that is, we begin with
\[ \delta E = \int d^3 r \delta \mathcal{H}(r) \]
\[ = \sum_{n' L' V' J' M'} (41) \]
An explicit variation over the primary densities under the differential operators \( D_{ml} \) can be avoided by first integrating by parts and recoupling. The recoupling within a scalar [5] is simple, namely,

\[ [A_{f}(B;C_{f})]_0 = (-1)^{l_f + l_c} [C_{f}(B;A_{f})]_0. \]

Hence, the integration by parts gives:

\[ \delta \mathcal{H}(r) = \sum_{n' L' V' J' M'} (-1)^{l_f + l_c} \left( \int d^3 r \left( \frac{\partial \rho_{n' L' V' J'}}{\partial \rho_{n L V J}} \right) \right) \delta \rho_{n' L' V' J' M'}. \]

where we have used Eq. (32), then we changed the names of indices, and we also used the fact that \( m + l \) is even.

Therefore, the variation under the differential operators \( D_{ml} \) only gives the transposition of indices \([n', L', V', J'] \leftrightarrow [nLJV] \) and the phase. The complete variation of the energy then reads:

\[ \delta E = \sum_{n' L' V' J'} \int d^3 r \left[ \delta \rho_{n' L' V' J'} \tilde{U}_{n' L' V' J'}(r) \right]_0. \]

where we defined potentials

\[ \tilde{U}_{n' L' V' J' M'}(r) = \sum_{mlnJV} \left( \sum_{m'} \left[ C_{mlnJV}^* C_{m'l'n'J'V'} \right] [D_{ml} \rho_{n L V J}(r)]_0 \right) \delta \rho_{n' L' V' J' M'}. \]

Note that because of the ordering (40), in Eq. (45) either the first or the second term is non-zero (for \([n', L', V', J'] \neq [nLJV] \)), or both terms add up to \( 2\delta \rho_{n' L' V' J'} \) (for \([n', L', V', J'] = [nLJV] \)).

3.3. Fields

Now we are in a position to perform the variation of densities over the wave functions. To this end, we assume that the non-local density matrices in Eq. (39) have the general form of

\[ \rho(r \sigma', r' \sigma') = \sum_{i} \phi_i(r \sigma) \psi_i(r' \sigma'). \]

This form allows us to carry out the variation for several important specific cases simultaneously. Namely, for the standard Hartree–Fock (HF) case one has:

\[ \psi_i(r' \sigma') = \phi_i^*(r' \sigma'), \]

and the sum runs over the occupied states only, \( i = 1, \ldots, A \). Similarly, for the BCS case, or for the Hartree–Fock–Bogoliubov case in the canonical basis, one has:

\[ \psi_i(r' \sigma') = v_i^2 \phi_i^*(r' \sigma'). \]

where \( v_i^2 \) are the occupation factors and the sum runs over the pairing window. For transition densities pertaining to the symmetry restoration, one has:

\[ \psi_i(r' \sigma') = \sum_{j} 0^{-1}_{ij}(\alpha) \phi_j^*(r' \sigma', \alpha). \]

where \( \phi_j(r' \sigma', \alpha) = R(\alpha) \phi_j(r' \sigma') \) are the wave functions transformed by the symmetry operator \( R(\alpha) \) and \( 0_{ij}(\alpha) \) is the overlap matrix:

\[ O_{ij}(\alpha) = \int d^3 r \sum_{\sigma} \phi_i^*(r \sigma, \alpha) \phi_j(r \sigma). \]

Finally, for the random phase approximation (RPA) amplitudes given by non-hermitian matrix \( \tilde{\rho}_{ij} \) one has

\[ \psi_i(r' \sigma') = \sum_{j} \tilde{\rho}_{ij}^* \phi_j^*(r' \sigma'). \]

To derive the fields, first we recall that the variation of the non-local density, over the wave function reads

\[ \delta \rho_{\eta \mu}(r, r') = \sum_{\sigma} \sum_{\sigma'} \frac{\partial \rho_{\eta \mu}(r, r')}{\partial \psi_i(r' \sigma')} \delta \psi_i(r' \sigma') \]

\[ = \sum_{\sigma} \sum_{\sigma'} \left( \phi_{\eta \mu}(r \sigma) \phi_i(r \sigma') \right) \delta \psi_i(r' \sigma'). \]

Therefore, variation of the primary density is given by

\[ \delta \rho_{\eta \mu}(r, r') = \sum_{\sigma} \sum_{\sigma'} \left( \phi_{\eta \mu}(r \sigma) \phi_i(r \sigma') \right) \delta \psi_i(r' \sigma'). \]

At this point, operators KLMV' mix derivatives acting on the variables \( r \) and \( r' \), cf. Eq. (18). By using the Wigner–Eckart theorem, we can express them as sums of products of derivatives \( D_{mlm'1} \) and \( D_{m'1m} \), which act on \( r \) and \( r' \), respectively, that is,\n
\[ K_{n' L' V' M'} = \sum_{m' l' v} K_{n' m' l' v}^{n L} \sum_{M'M'} C_{M'M'}^{V' M} D_{m'm} D_{m'l'}, \]

where the order of derivative is conserved,

\[ n' = m + m'; \]

and the triangle rule of angular momentum coupling must be obeyed. Numerical coefficients \( K_{n' m' l' v}^{n L} \) can be calculated by using methods of symbolic programming. At N3LO, only 91 coefficients \( K_{n' m' l' v}^{n L} \) are needed, so they can easily be precalculated and stored.

We now can insert expressions (53) and (54) into Eq. (44) and remove condition \( r' = r \) by adding the integral over \( r' \) and the function \( \delta(r - r') \), that is,

\[ \delta E = \sum_{n' L' V' J'} \int d^3 r \int d^3 r' \delta(r - r') \]

\[ \times \left( \sum_{\sigma} \sum_{\sigma'} \left[ \tilde{U}_{n' L' V' J'}(r) K_{n' L' V' \sigma \sigma'} \right] \right) \phi_i(r \sigma) \delta \psi_i(r' \sigma'). \]

This allows us to integrate by parts over \( r' \) and transfer the action of \( D_{m'm'} \) onto the delta function. With all the angular momenta couplings shown explicitly, this gives
\[ \delta E = \sum_{n'l'v'j'} \int d^3 r \int d^3 r' \sum_{m'i'M_i} \sum_{\sigma'\sigma} \{ D_{m'i'M_i}^T \delta(r - r') \} \]
\[ \times \sum_{\sigma'\sigma} \sum_{M'_{i',i}} \sum_{\mu'\mu} C_{M'_{i',i},\mu'\mu}^{M_{i'},\mu} \frac{(-1)^{j'-M'}}{\sqrt{2j'+1}} \bar{U}_{n'l'v'j',-M'}(r) \]
\[ \times k_{n'l'v'j'}^{n'i'i'} C_{M_{i'1},i'M_{i1}}^{M_{i},i} D_{m'i'M_i} \sigma_{\mu'\mu} \sigma_{\sigma'\sigma} \phi_i(r) \delta(r' r') \delta(r' r') \delta(r' r'). \]
(57)

The action of \( D_{m'i'M_i}^T \) onto the delta function can be replaced by that of \((-1)^m D_{m'i'M_i}^T\), and then the integration by parts over \( r \) gives
\[ \delta E = \sum_{n'l'v'j'} \int d^3 r \int d^3 r' \sum_{m'i'M_i} \sum_{\sigma'\sigma} \{ D_{m'i'M_i}^T \delta(r - r') \} \]
\[ \times \sum_{\sigma'\sigma} \sum_{M'_{i',i}} \sum_{\mu'\mu} C_{M'_{i',i},\mu'\mu}^{M_{i'},\mu} \frac{(-1)^{j'-M'}}{\sqrt{2j'+1}} \bar{U}_{n'l'v'j',-M'}(r) \]
\[ \times (-1)^m D_{m'i'M_i}^T \bar{U}_{n'l'v'j',-M'}(r) \]
\[ \times k_{n'l'v'j'}^{n'i'i'} C_{M_{i'1},i'M_{i1}}^{M_{i},i} D_{m'i'M_i} \sigma_{\mu'\mu} \sigma_{\sigma'\sigma} \phi_i(r) \delta(r' r') \].
(58)

In the resulting local integral we demand that the operator acting on \( \phi_i(r) \) is equal to the mean-field operator, which gives
\[
 h(\rho) = \sum_{n'l'v'j'} \sum_{m'i'M_i} \sum_{\sigma'\sigma} \sum_{\mu'\mu} C_{M'_{i',i},\mu'\mu}^{M_{i'},\mu} \frac{(-1)^{j'-M'}}{\sqrt{2j'+1}} \bar{U}_{n'l'v'j',-M'}(r) \]
\[ \times (-1)^m k_{n'l'v'j'}^{n'i'i'} C_{M_{i'1},i'M_{i1}}^{M_{i},i} D_{m'i'M_i} \sigma_{\mu'\mu} \sigma_{\sigma'\sigma} \phi_i(r) \delta(r' r'), \]
(59)

In this form, the mean-field operator is expressed through sums of derivative operators standing on both sides of the potentials \( \bar{U}_{n'l'v'j',-M'} \). This form can easily be used in the calculation of the matrix elements, because the left derivative operator can simply be applied onto the left wave function through the integration by parts. Moreover, potentials \( \bar{U}_{n'l'v'j',-M'} \) are related to the secondary densities by very simple relations (45). However, it turns out that numerical calculations are much faster if the derivatives appear only on one side of the potentials, like it was postulated in Eq. (7), that is,
\[
 h(\rho) = \sum_{n'l'v'j'} \left[ U_{n'l'v'j'}(r) D_{m'i'M_i} \sigma_{\mu'\mu} \right] \delta \rho. \]
(60)

3.4. Fields for terms containing additional density dependence

In the Skyrme functional, a dependence on the isoscalar density is usually added to the four terms that are zero order in derivatives, \( \rho_\sigma^2 \) and \( s_\sigma^2 \) \( (\tau = 0, 1) \), that is
\[
\rho_\sigma^2 = \sum_{\alpha=0,1} \left( \frac{\rho_\sigma^2}{\rho_\sigma^2} \right) \delta \rho, \]
(61)
\[
\rho_\sigma^2 = \sum_{\alpha=0,1} \left( \frac{\rho_\sigma^2}{\rho_\sigma^2} \right) \delta \rho. \]
(62)

We use the extra index \( \alpha \) on the coupling constant to distinguish the notation from the one used for the density-independent terms.

The contribution to the fields is obtained from the variation of the energy, which we show here only for the first of the above two terms, namely,
We now introduce coefficients $D_{\gamma\delta\rho\pi}^{(1)}$, which allow for expressing products of derivatives as:

$$D_{\gamma'\delta'\rho'\pi'}^{(1)} = \sum_{\mu\nu\mu'} C_{\gamma\mu\delta\mu'}^{\nu\mu} D_{\gamma'\delta'\rho\pi}^{(1)}$$

$$D_{\gamma'\delta'\rho'\pi'}^{(2)} = \sum_{\mu\nu\mu'} C_{\gamma'\delta'\rho'\pi'}^{\mu\nu} D_{\gamma\delta\rho\pi}^{(2)}$$

where $u = \gamma + \rho$ and $u' = \gamma' + \rho'$, that is, $u + u' = m + n$. These coefficients can be calculated by using methods outlined in Appendix A. At NLO, only 91 coefficients $D_{\gamma\delta\rho\pi}^{(1)}$ are needed, so they can easily be precalculated and stored.

The sum of products of four Clebsh–Gordan coefficients can now be recoupled (see Eq. 8.7[20] in Ref. [5]) as:

$$\sum_{M_1 M_2 M_3 M_4} c_{M_1 M_2 M_3 M_4}^{M_1 M_2 M_3 M_4} C_{M_1 M_2 M_3 M_4}^{M_1 M_2 M_3 M_4} = \sqrt{(2I + 1)(2L + 1)(2U + 1)}$$

$$\sum_{M_1 M_2 M_3 M_4} c_{M_1 M_2 M_3 M_4}^{M_1 M_2 M_3 M_4} C_{M_1 M_2 M_3 M_4}^{M_1 M_2 M_3 M_4} = \sqrt{(2I + 1)(2L + 1)(2U + 1)}$$

where $u = \gamma + \rho$ and $u' = \gamma' + \rho'$, that is, $u + u' = m + n$. These coefficients can be calculated by using methods outlined in Appendix A. At NLO, only 91 coefficients $D_{\gamma\delta\rho\pi}^{(1)}$ are needed, so they can easily be precalculated and stored.

The last two remaining Clebsh–Gordan coefficients can be absorbed in the following definition of the coupled derivative of the density:

$$D_{\gamma'\delta'\rho'\pi'}^{(2)}$$

$$\mu_\nu$$

$$\sqrt{(2W + 1)(2J + 1)[L I W] J}$$

where coefficients $A_{\mu_1 \mu_2 \nu_1 \nu_2}^{(1)}$ result from summing up all intrinsic indices:

$$A_{\mu_1 \mu_2 \nu_1 \nu_2}^{(1)} = \sum_{\mu_1 \mu_2 \nu_1 \nu_2} A_{\mu_1 \mu_2 \nu_1 \nu_2}^{(1)}$$

where $\gamma = \gamma + \rho$ and $\gamma' = \gamma' + \rho'$.

Finally, gives

$$\rho_{\mu_1 \mu_2 \nu_1 \nu_2}^{(1)} = \sum_{\mu_1 \mu_2 \nu_1 \nu_2} A_{\mu_1 \mu_2 \nu_1 \nu_2}^{(1)}$$

where coefficients $A_{\mu_1 \mu_2 \nu_1 \nu_2}^{(1)}$ are needed, so they can easily be precalculated and stored.

4. The N³LO potentials, fields, and densities in the spherical harmonic oscillator basis

4.1. Spherical HO basis

The standard spherical HO wave functions are given by

$$\phi_{\mu N}(r, \theta, \phi, \sigma)$$

$$= b^{3/2} F_{\mu \nu}(br) e^{-1/2(\sigma)} \sum_{m_\mu m_\pi} C_{m_\mu m_\pi}^{3m_\mu} \phi_{\mu N}(\theta, \phi) \chi_{2m_\pi}(\sigma).$$

where \(b\) is the oscillator constant.

$$b = \frac{\sqrt{\text{MO}}}{h}$$

and $F_{\mu \nu}(br)$ are proportional to the standard Laguerre polynomials [6].

To calculate the secondary densities $\rho_{\mu 1 \mu 2 \nu 1 \nu 2}$, one has to act on the space part of the wave function (78) with the derivative operators $D_{\mu 1 \mu 2}$. This leads to defining the polynomials $F_{\mu \nu}^{(1)}(br)$ such that

$$D_{\mu 1 \mu 2} b^{3/2} F_{\mu \nu}(br) e^{-1/2(\sigma)} \sum_{m_\mu m_\pi} C_{m_\mu m_\pi}^{3m_\mu} \phi_{\mu N}(\theta, \phi)$$

$$= b^{3/2} \sum_{m_\mu m_\pi} F_{\mu \nu}^{(1)}(br) e^{-1/2(\sigma)} \chi_{2m_\pi}(\sigma).$$

and our goal is to determine the set of reduced polynomials $F_{\mu \nu}^{(1)}(br)$, in terms of which the derivatives of the spherical HO wave functions read

$$D_{\mu 1 \mu 2} \phi_{\mu N}(r, \theta, \phi, \sigma)$$

$$= b^{3/2} \sum_{m_\mu m_\pi} \frac{1}{\sqrt{2k + 1}} F_{\mu \nu}^{(1)}(br) e^{-1/2(\sigma)} \chi_{2m_\pi}(\sigma).$$

Explicit form of $F_{\mu \nu}^{(1)}(br)$ can be calculated by using the Wigner–Eckart theorem again, namely, Eq. (80) must have the form

$$D_{\mu 1 \mu 2} \phi_{\mu N}(r, \theta, \phi, \sigma)$$

where $\phi_{\mu N}$ is the space part of the wave function (78). Then we have polynomials $F_{\mu \nu}^{(1)}(br)$ expressed through the Laguerre polynomials as:

$$D_{\mu 1 \mu 2} \phi_{\mu N}(br) = D_{\mu 1 \mu 2} \phi_{\mu N}(br) = b^{3/2} F_{\mu \nu}^{(1)}(br) = \sum_{N'} (\phi_{N'K} || D_{\mu 1 \mu 2} || \phi_{\mu N}) F_{N'K}(br).$$

where the reduced matrix element can be calculated by considering only one matrix element, namely.
\begin{align}
\langle \phi_{N'k,m_0=0}|D_{nl,M_0}=0|\phi_{NI,m}=0 \rangle \\
= \frac{1}{\sqrt{2k+1}} C_{l0}^{l0}(\phi_{N'k}|D_{nl}|\phi_{NI}).
\end{align}

Note that the Clebsch–Gordan coefficient $C_{l0}^{l0}$ is not zero for angular momenta restricted by the parity conservation, $(-1)^{l+l'-k} = 1$.

The parity conservation induces specific conditions on the polynomials $F_{nLNI}(br)$. Indeed, by comparing parities of both sides of Eq. (80) we see that
\begin{equation}
F_{nLNI}(br) = 0, \quad \text{for } (-1)^{n+l+k} = -1.
\end{equation}

Equivalently, since for all derivative operators we have $(-1)^{n+l} = 1$, we see that
\begin{equation}
F_{nLNI}(br) = 0, \quad \text{for } (-1)^{l+l+k} = -1.
\end{equation}

Since polynomials $F_{nLNI}(br)$ are real, phases of polynomials $F_{nLNI}(br)$ are fixed by those of the derivative operators \((30)\) and spherical harmonics \([5]\),
\begin{equation}
Y_{J'M'}^\ast(\theta, \phi) = (-1)^M Y_{J'-M}^\ast(\theta, \phi),
\end{equation}
that is,
\begin{equation}
F_{nLNI}(br) = (-1)^{l-k} F_{nLNI}(br) = (-1)^n F_{nLNI}(br)
\end{equation}
where the last two equivalent forms result from Eqs. (86) and (87).

The $F_{nLNI}$ polynomials are calculated using formulas for spherical derivatives (see \([5]\)) combined with recursion relations for derivatives of Laguerre polynomials. In this way spherical derivatives of one of the basis functions $\phi_{nLNI}(r, \sigma) = g_{IL}(r) C_{l0}^{l0} Y_{lm}(\theta, \phi)$ can be expressed as a sum of functions
\begin{equation}
\nabla_{I_1} \phi_{nLNI} \cdots \nabla_{I_k} \phi_{nLNI} Y_{lm}(\theta, \phi)
\end{equation}
\begin{equation}
= \sum_{l_1, \ldots, l_N = -1} \left[ a(l, n, I_1, \ldots, I_N, r) g_{n-l_1, l}(r) \\
+ b(l, n, I_1, \ldots, I_N, r) g_{n-l_1, l}(r) \right] Y_{l_1+l_1+\cdots+l_N+m+\mu_1+\cdots+\mu_N},
\end{equation}
where the $a$ and $b$ coefficients needed for the different orders were derived using symbolic programming. In this way we obtain analytical expressions for all derivatives which can then be calculated with good accuracy.

\subsection{Densities in the spherical HO basis}

For any density matrix one can always perform a multipole expansion. This strategy fits very well our applications to the spherical HF solutions, where only the monopole component of the density matrix is non-zero, and to the RPA applications in spherical nuclei, where all multipole excitations separate from one another. Therefore, in what follows we consider the density matrix of multipolarity $J$ in the form given by the Wigner–Eckart theorem:
\begin{equation}
\hat{\rho}_{nLNI,N'T'j'm'} = \frac{1}{\sqrt{2J+1}} C_{j'm'|JM}^{JM} \langle \phi_{nLNI}|\hat{\rho}_J|\phi_{N'T'j'm'} \rangle,
\end{equation}
and depending on its reduced matrix elements ($\phi_{nLNI}|\hat{\rho}_J|\phi_{N'T'j'm'}$), Then, the non-local densities can be expressed in terms of the spherical HO wave functions \((78)\) as
\begin{align}
\hat{\rho}_{\nu\mu}(r_1, r_2) &= \sum_{N'T'j'm'} \phi_{N'T'j'm'}(r_1, \sigma) \hat{\rho}_{nLNI,N'T'j'm'} \times \phi_{N'T'j'm'}(r_2, \sigma') \langle \sigma' | \phi_{\nu\mu} | \sigma \rangle,
\end{align}
that is,
\begin{equation}
\hat{\rho}_{\nu\mu}(r_1, r_2) = \sum_{N'T'j'm'} \phi_{N'T'j'm'}(r_1, \sigma) \hat{\rho}_{nLNI,N'T'j'm'} \times \phi_{N'T'j'm'}(r_2, \sigma') \langle \sigma' | \phi_{\nu\mu} | \sigma \rangle.
\end{equation}

We can now replace the spin coordinates by the spin projections,
\begin{equation}
\chi_{2m}(\sigma) = \delta_{m, \sigma}, \quad \chi_{2m}(\sigma) = \delta_{m, \sigma},
\end{equation}
and we can use the Wigner–Eckart theorem for the Pauli matrices,
\begin{equation}
\langle \frac{1}{2} | \sigma \rangle \langle \frac{1}{2} | \frac{1}{2} = -i \sqrt{6},
\end{equation}
After inserting the non-local density \((92)\) into the expression for local densities \((75)\), and after acting with derivatives on spherical wave functions, as in Eqs. (80) and (81), we obtain
\begin{align}
\hat{\rho}_{\nu\mu}(r) &= \sum_{M\mu} C_{W\mu}^{W\mu}(r) \sum_{M\mu} C_{W\mu}^{W\mu}(r) \langle \phi_{nLNI}|\hat{\rho}_J|\phi_{N'T'j'm'} \rangle \times \sum_{N'T'j'm'} \phi_{N'T'j'm'}(r_1, \sigma) \hat{\rho}_{nLNI,N'T'j'm'} \times \phi_{N'T'j'm'}(r_2, \sigma') \langle \sigma' | \phi_{\nu\mu} | \sigma \rangle,
\end{align}

where, by using the phase convention \((2)\)
\begin{equation}
\hat{\rho}_{\nu\mu}(r) = \sum_{\nu\mu} \phi_{\nu\mu}(r_1, \sigma) \phi_{\nu\mu}(r_2, \sigma') \langle \sigma' | \phi_{\nu\mu} | \sigma \rangle.
\end{equation}

for the radial form factors \((79)\) given by:
\[ \mathcal{F}^{\mu U'u'W',J'}_{\nu v J} (br) = (-1)^{1+\nu} \left( \begin{array}{c} 1 \\ \frac{1}{2} \end{array} \right) \left( \begin{array}{c} \sigma_{\nu} \\ \frac{1}{2} \end{array} \right) \sqrt{\frac{2J'+1}{4\pi}} \times \sum_{Nlj} \mathcal{P}^{j+k W'J'}_{k+k U',v J} \rho_{k+k 0}^{j+k W'J'} \times F^k_{U'Nlj} (br) (|\phi_{Nlj}| |\phi_{Nlj}|) F^{\nu U'W'}_{\mu U'W'} (br), \]

(98)

where

\[ \mathcal{P}^{j+k W'J'}_{k+k U',v J} = (-1)^j (-1)^T (-1)^{U'}C^{T0}_{k+k 0} \times \left( \frac{(2k+1)(2k'+1)(2W+1)(2j+1)(2J'+1)}{(2T+1)} \right) \times \sum_{T'} (-1)^T (2T'+1) \times \left\{ \begin{array}{c} j' j' J'' \\ \frac{1}{2} \frac{1}{2} v \\ \frac{1}{2} \frac{1}{2} v' \end{array} \right\} \left\{ \begin{array}{c} I' I' T' \\ k' k' U' \\ T' T' W \end{array} \right\} \left\{ \begin{array}{c} v T' J'' \\ U' W' \end{array} \right\} \right\}. \]

(99)

In view of the fact that only the coefficients \( \mathcal{P}^{j+k W'J'}_{k+k U',v J} \) for \((-1)^{j+k W'U} = 1 \) and \((-1)^{j+k W'U'} = 1 \) are required in Eq. (98), we may replace them by coefficients \( \bar{\mathcal{P}}^{j+k W'J'}_{k+k U',v J} \):

\[ \bar{\mathcal{P}}^{j+k W'J'}_{k+k U',v J} = (-1)^j C^{T0}_{k+k 0} \sqrt{\frac{(2k+1)(2k'+1)(2W+1)(2j+1)(2J'+1)}{(2T+1)}}, \]

(100)

where we have used symmetry properties of \( J \) symbols under the transposition of rows and columns and transposition with respect to the main diagonal.

Finally, all secondary densities of Eq. (76) can now be calculated in terms of one compact expression:

\[ \mathcal{F}^{J'M'R}_{\mu nlv_{\nu} J'M'R} (br) = \sum_{TM_{\nu}} \mathcal{P}^{J'M'R}_{\mu nlv_{\nu} J'} (br) e^{-b(r)^3} T^{J'M'R}_{TM_{\nu}} Y_{TM_{\nu}} (\theta, \phi), \]

(101)

where the radial form factors read

\[ \mathcal{P}^{J'M'R}_{\mu nlv_{\nu} J'} (br) = \sum_{U'U'W'W'} A^{\mu U'u'U'W'}_{\nu v W'} T^{U'u'U'W',J'}_{\mu U'u'U'W',J'} (br). \]

(102)

4.3. Potentials in the spherical HO basis

In the spherical basis, the secondary densities to be used in Eq. (45) have the form given in Eq. (101). Therefore, the potentials in Eq. (59) acquire the form

\[ \bar{U}^{J'M'R}_{W'U'J'M'R} (r) = \sum_{TM_{\nu}} \bar{U}^{J'M'R}_{W'U'J'} (br) e^{-b(r)^3} T^{J'M'R}_{TM_{\nu}} Y_{TM_{\nu}} (\theta, \phi), \]

(103)

where the radial form factors read:

\[ \bar{U}^{J'M'R}_{W'U'J'} (br) = \sum_{TM_{\nu}} \left\{ C^{J'M'R}_{\mu nlv_{\nu} J'} (br) + (-1)^{J'-J} C^{J'M'R}_{\mu nlv_{\nu} J'} (br) \right\} \rho_{\mu nlv_{\nu} J'} (br). \]

(104)

Similarly, potentials \( U^{J'M'}_{W'U'J'M'} (r) \) in Eq. (60) read

\[ U^{J'M'}_{W'U'J'M'} (r) = \sum_{TM_{T}} U^{J'M'}_{W'U'J'} (br) e^{-b(r)^3} T^{J'M'}_{TM_{T}} Y_{TM_{T}} (\theta, \phi), \]

(105)

and the radial from factors \( U^{J'M'}_{W'U'J'} (br) \) can be calculated in the complete analogy to the results outlined in Section 2, see Eq. (8), namely

\[ U^{J'M'}_{W'U'J'} (br) = \sum_{M_{T}} C^{\beta \alpha}_{\alpha \beta} \rho_{\alpha \beta M_{T} M_{T}'} \rho_{\mu nlv_{\nu} J'} (br). \]

(106)

4.4. Matrix elements of the Hamiltonian (59) in the spherical HO basis

In the spherical HO basis of Eq. (78), we can calculate the matrix elements of the single-particle Hamiltonian (59) in the following way:

\[ \langle N'J'J'm_j' | \bar{h} | Nljm_j \rangle = \sum_{n'U'V'J'M'R} \sum_{M_{T}} \sum_{M_{T}'} \int r^2 dr \int d\Omega \]

\[ \times \sum_{\sigma' \sigma M'} \sum_{\sigma' \sigma M'} C^{J'M'R}_{U'V'M'} (br) e^{-b(r)^3} T^{J'M'R}_{TM_{T}'} Y_{TM_{T}'} (\theta, \phi), \]

(107)

where \( \bar{h}(\rho^{J'M'}) \) denotes the field calculated for the one-multipole density matrix \( \rho^{J'M'} \) (90). The integration by parts now gives

\[ \langle N'J'J'm_j' | \bar{h} | Nljm_j \rangle = \sum_{n'U'V'J'M'R} \sum_{M_{T}} \sum_{M_{T}'} \int r^2 dr \int d\Omega \]

\[ \times \sum_{\sigma' \sigma M'} \sum_{\sigma' \sigma M'} \bar{C}^{J'M'R}_{U'V'M'} (br) e^{-b(r)^3} T^{J'M'R}_{TM_{T}'} Y_{TM_{T}'} (\theta, \phi), \]

(108)

where we have used the hermitian-conjugation property (31) of the differential operators.

By inserting potentials (103) and derivatives of spherical wave functions (82) we have:

\[ \langle N'J'J'm_j' | \bar{h} | Nljm_j \rangle = \sum_{n'U'V'J'M'R} \sum_{M_{T}} \sum_{M_{T}'} \int r^2 dr \int d\Omega \]

\[ \times \sum_{\sigma' \sigma M'} \sum_{\sigma' \sigma M'} \bar{C}^{J'M'R}_{U'V'M'} (br) e^{-b(r)^3} T^{J'M'R}_{TM_{T}'} Y_{TM_{T}'} (\theta, \phi), \]

(109)
\[ \times C_{1M1'F M_1}'^L (\sigma') (\sigma'')(\sigma) (-1)^{m+1-M} \]

\[ \times \sum_{TM_1'} U_{n' l' v' j'} (br) e^{-ibr^2} C_{1M1'F M_1}'^L Y_{TM_1} (\theta, \phi) \]

\[ \times b^{m+3/2} \sum_{k_m k_{m_1}} \frac{1}{\sqrt{2k + 1}} C_{k_m}^k C_{m}^m C_{m_1}^m \]

\[ \times F_{m_1 M_1 N_l}^k (br) e^{-\imath br} Y_{k_m} (\theta, \phi) \chi_{1 m_1}^m (\sigma') \]

\[ \times b^{m+3/2} \sum_{k_m k_{m_1}} \frac{1}{\sqrt{2k + 1}} C_{k_m}^k C_{m}^m C_{m_1}^m \]

\[ \times r^{k}_{m_1 N_l} (br) e^{-\imath br} Y_{k_m} (\theta, \phi) \chi_{1 m_1}^m (\sigma'). \]  

The angular part can be integrated explicitly, by using the multiplicity law of spherical harmonics (Eq. 5.6(9) in Ref. [5]), and summations over \( \sigma \) and \( \sigma' \) can be performed as in Eqs. (93)–(95). This gives

\[ \langle N'I' j'm' | \tilde{h} (\tilde{\rho} J'M') | Nljm \rangle = \sum_{n' l' v' j'} \sum_{M' l' M_1'} \sum_{M_1 M} \int r^2 dr \]

\[ \times C_{n M_1 M_1' M}^{l' M_1' J'M'} (-1)^{m+1-M} \sum_{T M_1} C_{n M_1 M_1' M}^{l M_1 J M_1'} \]

\[ \times (-1)^{m'+1-M'} \sum_{T M_1} \sum_{m} \sum_{m_1} \sum_{m_1'} \frac{1}{\sqrt{2j + 1}} C_{m}^m C_{m_1}^{m_1} \]

\[ \times \left( \frac{1}{2} \right | | \sigma_v' \rangle \left( \frac{1}{2} \right | | \sigma_v \rangle \right) \sum_{T M_1} \]

\[ \times F_{m_1 N_l}^k (br) U_{n' l' v' j'} (br) F_{m_1 M_1}^k (br) e^{-2ibr^2} \]

\[ \times C_{T M_1}^{T M_1} \]

\[ \times \frac{1}{\sqrt{2k + 1}} C_{k_m}^k C_{m_1}^{m_1} C_{m}^m. \]  

where we have used condition (55).

We may now proceed by calculating the reduced matrix element of the field:

\[ \langle \phi_{N'T'} j' | \tilde{h} (\tilde{\rho} J'M') | \phi_{Nljm} \rangle = \sum_{m_1 M_1} \frac{1}{\sqrt{2j + 1}} C_{m_1}^{m_1} \langle N'I' j'm' | \tilde{h} (\tilde{\rho} J'M') | Nljm \rangle. \]  

Again, after a lengthy but straightforward derivation presented in Appendix C, we obtain the following result:

\[ \langle \phi_{N'T'} j' | \tilde{h} (\tilde{\rho} J'M') | \phi_{Nljm} \rangle = \delta^{j'-j} \sum_{n' l' v' j'} \sum_{M' l' M_1'} \sum_{M_1 M} \]

\[ \times \left( \frac{1}{2} \right | | \sigma_v \rangle \left( \frac{1}{2} \right | | \sigma_v' \rangle \right) \sum_{T M_1} \]

\[ \times F_{m_1 N_l}^k (br) U_{n' l' v' j'} (br) F_{m_1 M_1}^k (br) e^{-2ibr^2} \]

\[ \times C_{T M_1}^{T M_1} \]  

where we see the same numerical coefficients (99) that already appeared in Eq. (98).
which, for density-independent coupling constants, can be integrated over \( \theta \) and \( \phi \). After summing up the Clebsh–Gordan coefficients, one obtains:

\[
\xi_{\ell m \ell' m'}^{\alpha \beta} = \frac{(-1)^{\ell - M_0} \delta_{J_0}^{J_0} \delta_{M_0'}^{M_0} |2T + 1|} {2 \sqrt{T + 1}} \int r^2 dr C_{\ell m \ell' \ell'}^{\alpha \beta} e^{-2br^2} \sum_T (-1)^T (2T + 1) \\
\times \hat{\rho}_0^{J_0} (\beta) \hat{T}_0^{J_0} (\alpha) \\
\times \sum_{M_0} Q_{M_0} \left\{ \begin{array}{ccc} J_0' & J_0' & 0 \\ M_0' & J_0' & M_0 \\ J_0' & J_0' & Q \end{array} \right\} Y_Q M_0 (\theta, \phi).
\]

One thus obtains the correct result that for density-independent coupling constants, the potential energy is diagonal in different multipoles.

When the coupling constants do depend on density, the angular integral cannot be performed, and a similar derivation gives the general result:

\[
\xi_{\ell m \ell' m'}^{\alpha \beta} = \int d^3 r \ C_{\ell m \ell' \ell'}^{\alpha \beta} e^{-2br^2} \\
\times \hat{\rho}_0^{J_0} (\beta) \hat{T}_0^{J_0} (\alpha) \\
\times \sum_{M_0} Q_{M_0} \left\{ \begin{array}{ccc} J_0' & J_0' & 0 \\ M_0' & J_0' & M_0 \\ J_0' & J_0' & Q \end{array} \right\} Y_Q M_0 (\theta, \phi).
\]

For the spherical case, one has \( J_0' = J_0 = 0 \), which implies \( T = J' \), and both expressions (117) and (118) reduce to:

\[
\xi_{\ell m \ell' m'}^{\alpha \beta} = ( -1 )^{J_0'} \sqrt{2J_0' + 1} \int r^2 dr C_{\ell m \ell' \ell'}^{\alpha \beta} e^{-2br^2} \\
\times \hat{\rho}_0^{J_0} (\beta) \hat{T}_0^{J_0} (\alpha).
\]

For density-dependent terms, such as those in Eqs. (61) and (62), the total energy reads:

\[
\xi_{\ell m \ell' m'}^{\alpha \beta} = ( -1 )^{J_0'} \sqrt{2J_0' + 1} e^{\alpha \ell m \ell' \ell'} f ( \rho_0 ) \int r^2 dr e^{-2br^2} \rho_0 (br) \\
\times \hat{\rho}_0^{J_0} (\beta) \hat{T}_0^{J_0} (\alpha) \hat{\rho}_0^{J_0} (\alpha).
\]

where the density \( \rho_0 (br) \) depends on the reduced density as:

\[
\rho_0 (br) = \rho_{0,000,000}(br) e^{-(br)^2} / \sqrt{4\pi}.
\]

4.7. Direct Coulomb energy in the spherical symmetry

The integral for the direct Coulomb energy in spherical symmetry has a integrand which has a discontinuous first derivative. In order to get an integral which is easier to calculate using Gaussian–Hermite integration we use the ‘Vautherin trick’ [7]

\[
\frac{e^2}{2} \int d^3 r d^3 r' \rho (r) \frac{1}{|r - r'|} \rho (r') \\
= \frac{(4\pi)^2 e^2}{6} \int dr dr' (r + r')^3 - |r - r'|^3 |r' \rho (r') \Delta' \rho (r') \\
= \frac{4\pi e^2}{2\sqrt{3}} \int dr dr' \hat{\rho}_{0,000,000}(r)r e^{-(br)^2} \\
\times \int dr (r + r')^3 - |r - r'|^3 |r' \rho_{0,000,000}(r') e^{-(br)^2} \\
= \frac{4\pi e^2}{2\sqrt{3}} \int \hat{\rho}_{0,000,000}(r)r e^{-(br)^2} V_D (r).
\]

which gives a smoother integrand. However in order to perform the integral–response calculations, where expressions for non-spherical multipoles are needed, we also developed a different method which will be presented elsewhere [8].

4.8. Exchange Coulomb energy

For the exchange Coulomb energy we use the Slater approximation

\[
E_{ex} = -\frac{3}{4} \left( \frac{3}{\pi} \right)^{1/3} e^2 \int dr \rho_p^{A/3} \\
= -\frac{3}{4} \left( \frac{3}{\pi} \right)^{1/3} e^2 4\pi \int r^2 dr \hat{\rho}_{p,000,000,0}^{A/3} e^{-\frac{1}{4}(br)^2} \\
= -\frac{3}{4} \left( \frac{3}{\pi} \right)^{1/3} e^2 \int \left[ \hat{\rho}_{p,000,000,0}^{A/3} e^{-\frac{1}{4}(br)^2} \right] e^{-2(b'br)^2} r^2 dr.
\]

The matrix elements are the same as with density-dependent forces (using \( \alpha = -2/3 \) and \( e^{-\alpha \rho_0} = -\pi^{1/(\pi^2 e^2) \beta} \)) but should only be added to the proton matrix elements.

4.9. Numerical integration

Numerical Gauss–Hermite integration is used to calculate the radial integrals occurring in the expressions for matrix elements [Eqs. (112) and (113)] and total energies [Eq. (119)]. This kind of integration is for integrals of the form \( \int_0^\infty e^{-x^2} f(x) \, dx \), and in order to obtain this form our integrals are transformed by using \( x' = (\sqrt{2b}) x \). The integrals can then be written as:

\[
\int_0^\infty e^{-2(bx)^2} f (bx) \, x^2 \, dx \\
= \frac{1}{2} \int_{-\infty}^{\infty} e^{-x^2} \int_{-\sqrt{2b} X}^{\sqrt{2b} X} f (x^2 X) \, dx' \\
= \frac{1}{2} \int_{-\infty}^{\infty} e^{-x^2} \int_{-\sqrt{2b} X}^{\sqrt{2b} X} f (x^2 X) \, dx' \\
\approx \frac{1}{\sqrt{2b^3}} \sum_{i=1}^{N/2} w_i f (x_i^2) \, (x_i^2)^2.
\]

where \( f_{sym}(x) = f(x) \theta(x) + f(-x) \theta(-x) \) was used in an intermediate step. To reduce the number of grid points by half to \( N_{grid} = N/2 \) it was also used that for Gaussian–Hermite integration, the weight functions \( w_i \) and grid points \( x_i^2 = X_i^2 = br_i \) are symmetric around the origin. The integrals for matrix elements and total energies of most terms become exact when \( N_{grid} = N_0 + 2 \), where \( N_0 \) denotes the maximum HO shell included in the basis. But in general more points are needed when the integrand cannot be expressed as a product of four basis states, e.g., in the case for the Coulomb interaction and also for the density-dependent terms.

5. Overview of the software structure

In order to get an idea of the software structure we list some of the main calls performed by the program. In the code HOSPH (v1.02), there are also deeper level calls that are not listed here. Depending on the input data, some of the calls listed below may or may not be performed, but all are anyway included in the list. We specify the subroutine calls by first writing the name of the module followed by the name of the subroutine.
6. Description of the individual software components

The code HOSPHE (v1.02) is separated into different files with each file containing a module. Modules may contain module parameters and collections of subroutines and functions. Below follows a list of the files defining the code where the *’s should be replaced by the latest version number.

The modules which are used mainly in the initialization phase are:

- **hfmain**
  Driver program which reads input and starts the run by calling the hf routine.
  - **hf**
    Main program which contains the iteration loop. Starts by calling routines which initializes different tables.
    * RecoupDK: define_cfk
    * RecoupDK: define_cfd
    * RecoupA: define_cfa
    * RecoupA: define_cfa2
    * Coupling: restrict_densities
    * Coupling: get_cc_restrict
    * Coupling: make_ltab
    * Skyrme: get_skyrme_cc
    * Coupling: read_ccs_from_file
    * Coupling: cc_perl_to_n3lo
    * Coupling: set_dependent_cc
    * Coupling: cc_n3lo_to_perl
    * grid: set_grid
    * qnotra: set_sps
    * qnotra: qbase
    * mflib: Read_hf_initial_density_matrix
    * mflib: set_onedet
    * Ho_Derivatives: store_derivatives_of_basis
    * Ho_Derivatives: read_derivatives_of_basis
    * TypeDefinitions: inverse_derivative_densities
    * F2U: Define_F2U
    * mflib: read_density_matrix

Next follows the main iteration loop.

- **Derivative_densities**: sum_derivative_densities
- **hmatrix_Ud**: hamiltonian_matrix
- **mflib**: diagonalize_hf
- **mflib**: hf2ho
- **density_energy**: energy_from_densities
- **density_energy**: energy_from_densities_densdep
- **density_energy**: energy_from_coulomb
- **mflib**: hf_energy

Finally there is some post-processing of the results.

- **density_energy**: plot_energy_terms2
- **sorting**: sortx
- **mflib**: store_density_matrix

The modules which are used mainly in the iteration loop are:

- **hf.v*.f90**
  The main subroutine containing the iteration loop.

- **Derivative_densities.v*.f90**
  Generates the reduced derivative densities as functions of the grid points given a reduced density matrix as input.

- **hmatrix_Ud.v*.f90**
  Determines the matrix elements by using for the fields the form in Eq. (7). Calculates the reduced matrix elements given reduced densities and reduced matrix element quantum numbers as input. Also calculates the matrix elements of the density-dependent terms and of the Coulomb interaction.

- **matrix.v*.f90**
  Determines the matrix elements by using for the fields the form in Eq. (59). Calculates reduced matrix elements given reduced densities and reduced matrix element quantum numbers as input. Used mainly to test the accuracy of the matrix elements from hmatrix_Ud.v*.f90.

- **density_energy.v*.f90**
  Calculates integrals by using the Gauss–Hermite integration. Contains subroutines to calculate integrals over the terms in the functional to obtain the energy. Also calculates Coulomb energy integrals and radii.

- **coulomb.v*.f90**
  Calculates Coulomb direct and exchange matrix elements. Used mainly to test the accuracy of the matrix elements from hmatrix_Ud.v*.f90.

- **densitydep.v*.f90**
  Contains density-dependent factors that multiply each density-dependent coupling constants in the mean field potentials.

There are also modules containing helper routines:

- **sorting.v*.f90**
  Contains routines used to sort data.

- **geometric.v*.f90**
  Contains a collection of routines for the Clebsh–Gordan coefficients, 3, 6, and 9j symbols, Gamma functions, etc.

- **mflib.v*.f90**
  Contains a collection of auxiliary routines such as the diagonalization routines and various transformations.
7. Differences between the notation in the code HOSPHE (v1.02) and in the text

In some cases, the notation used in this study differs from that used in the code HOSPHE (v1.02). For example, to store real objects in the input, we have removed complex phases from some coefficients. The transformation rules for some useful objects are defined in Table 1.

8. Description of input data

Input is given to the code HOSPHE (v1.02) by using the FORTRAN “namelist” statement. In this way, the variables specified in the input have their values assigned, while those not specified in the input retain their predefined default values. The variables that can be specified in the input are listed below.

1. nmax
   Maximum main HO quantum number N0 used in the HO basis. The code HOSPHE (v1.02) currently supports values of N0 up to 70.

2. ordermax
   Maximum order in derivatives used. The possible choices are 0, 2, 4 and 6.

3. grid
   Number of Gauss–Hermite grid points. The code HOSPHE (v1.02) currently supports up to 85 grid points. If a negative value is given, the code uses N grid = N0 + 2 + 10 points, that is, 10 more points than is needed for the most terms to become exact. However, for the Coulomb and density-dependent terms to converge with high precision, one may need more grid points.

4. intera
   Name of the Skyrme functional to be used. At present, supported versions are: “SLY4” [9], “SLY4b” [10], “SLY5” [9], “SKM” [11], “SKP” [12], “SIII” [13], and “FILE”. If the name “FILE” is specified, the coupling constants are read from file “cc.inp”. Two example files of this type are included in the distribution. It is possible to either specify the coupling constants in the spherical notation as shown in file cc.inp_spherical or alternatively for a second order functional one may specify them using an alternative notation as illustrated in file cc.inp_cartesian.

5. AZ, AN
   Number of protons and neutrons.

6. hbarom
   Oscillator frequency \( \hbar \omega \) in MeV. If a negative value is given, it is calculated as \( \hbar \omega = 1.2 \times 41.4^{1/3} \), where A is the number of nucleons (see above).

7. boscil
   Oscillator constant \( b \) in \( \text{fm}^{-1} \). If a negative value is given, it is calculated from Eq. (79) by using the value of \( \hbar \omega \) (see above) and the nucleon mass \( m \) being the average of the neutron and proton masses.

8. icm
   Center of mass correction. For \( \text{icm} = 0 \), no correction is used, and for \( \text{icm} = 1 \), the code HOSPHE (v1.02) uses the one-body center of mass correction \( E_{\text{cm}} \approx -\frac{1}{A}(T) \).

9. icoudir
   For \( \text{icoudir} = 0 \), no direct Coulomb term is included, and for \( \text{icoudir} = -1 \), the code HOSPHE (v1.02) calculates the direct Coulomb energy by using the Vautherin method, see Section 4.7.

10. icouex
    For \( \text{icouex} = 0 \), the Coulomb exchange term is not included, and for \( \text{icouex} = -1 \), the code HOSPHE (v1.02) calculates the Coulomb exchange energy by using the Slater approximation, see Section 4.8.

11. itermx
    Maximum number of iterations allowed before aborting.

12. epsilon
    Accuracy parameter. The iterations are stopped when the ground-state energies calculated by using the EDF and HF expressions differ by less than epsilon and every HF single-particle energy changes less than epsilon between two iterations.

13. alpha
    Mixing parameter \( \alpha \) to slow-down/accelerate the iteration convergence. It mixes the density matrix from the current \( \rho_1 \) and previous \( \rho_0 \) iterations, so that the new density matrix is obtained as \( \rho = \alpha \rho_1 + (1 - \alpha) \rho_0 \).

14. keta J
    Turns the Skyrme tensor coupling constants ON/OFF (1/0).

15. restart
    If 1 or 2, the code HOSPHE (v1.02) attempts to read the density matrix from the file named as in the following example: densities_050_082.rec for \( (\text{AZ}, \text{AN}) = (50, 82) \) and the density matrix is also automatically stored to the same file when the iterations are finished. For restart = 0 no restart of iterations is attempted. For restart = 2 the stored density matrix is used even though it may come from a calculation with a different number of oscillator shells. For restart = 1 it is used only if the number of oscillator shells is the same.

16. Flag_read_ini_dm
    If .true., the code HOSPHE (v1.02) attempts to read the initially occupied levels from the file “occ_orbs.inp”. The first line should say \( N_{\text{eng}} \), \( N_{\text{fib}} \) where all main oscillator shells up to \( N_{\text{fib}} \) are automatically filled and \( N_{\text{eng}} \) denotes the number of changes in occupation with respect to this initial filling. Then follows one line per change, each change being specified as \( N, I, 2j, I_{\text{occ}} \) where \( N, I, 2j \) denotes quantum numbers of HO j-shells and \( I_{\text{occ}} \) specifies weather the shell should be occupied \( (I_{\text{occ}} = 1) \) or not \( (I_{\text{occ}} = 0) \). This is repeated twice to define shell occupancies for protons and then for neutrons. An example file of this type is included in the distribution.

17. verbose
    Verbose is an integer which specifies the amount of output produced by the code HOSPHE (v1.02) during a run. Verbose = 0 is the standard which gives a minimum of output and higher values leads to more information being printed to the screen.

9. Installation instructions

Most of the source code of HOSPHE (v1.02) is written in Fortran 95. The code consists of several modules that are linked together by using the standard Makefile (see [14]). In general, no special compiler options are needed, except for the optimization flags or check flags for tests. These options are set in the accompanying...
the least-square fit of the limiting values formed for $^{208}$Pb and the HO bases of main section of this file, which gives the total energies in $^{208}$Pb, which is also provided in the distribution file. The hosphe (v1.02). By executing the script, one obtains the output file "pb208.n50.out", which is also provided in the distribution file. The main section of this file, which gives the total energies in $^{208}$Pb calculated for the maximum HO shell included in the basis of $N_0 = 50$ and SLy4b Skyrme functional [10], reads

```plaintext
./hosphe << end > pb208.n50.out
&input AN=126,AZ=82,noscmx=50,
icm=1,icoudir=-1,icouex=-1,keta_J=1,
tera="SLy4b",ormax=2,
epsilon=1.e-7,itermax=1000,
Flag_read_ini_dm = .false.,restart = 0,
alpha=0.65,ngrid=-80,boscil=-2,hbarom=-1/
end
```

The script above is provided in the distribution file of the code hosphe (v1.02). By executing the script, one obtains the output file "pb208.n50.out", which is also provided in the distribution file. The main section of this file, which gives the total energies in $^{208}$Pb calculated for the maximum HO shell included in the basis of $N_0 = 50$ and SLy4b Skyrme functional [10], reads

```
Kin.prot   Kin. neut.   Tot. kin.
1337.059947 2529.116266 3866.176214

T=0 Skyrme         T=1 Skyrme         Tot. Skyrme
-6405.081099   106.598348   -6298.482751

Energy          HF Energy          Rearr. ene
-1635.692396   -1635.692396   -1221.821085

796.614142   827.882912   -31.268770
```

In Figs. 1 and 2, we present results of similar calculations performed for $^{208}$Pb and the HO bases of $N_0 = 10$–70. Fig. 1 shows the convergence of the total energy in function of $N_0$. It turns out that the energy converges exponentially to the limiting value of $E_0$, namely,

$$ E(N_0) = E_0 + \alpha N_0 \exp(-\alpha N_0). \quad (123) $$

However, as shown in the two panels of Fig. 1, two different values of $E_0$ and $\alpha$ are obtained for the regions of $N_0$ below and above $N_0 = 38$. A rather rapid convergence ($\alpha = 0.172(3)$) to $E_0 = -1635.719$, which is seen below $N_0 = 38$, is followed by a slower convergence ($\alpha = 0.1068(8)$) to $E_0 = -1635.69405$. Since the least-square fit of the limiting values $E_0$ is ill-conditioned, no error estimates can be obtained for them.

By considering convergence patterns in a few more cases for different options and nuclei, we found that the trend with two different slopes is not a general feature. In the few cases we looked at, we found that it is only above 40–50 shells that the rate seems to stabilize to an exponential convergence. These results show that, in general, it is not possible to find the extrapolated limit of energy just by calculating only a few points of the curve for some small numbers of shells.

Fig. 2 shows the dependence of CPU times on $N_0$, obtained on the AMD Opteron Processor 2352 running at 2100 MHz clock speed. First, one can see that the spherical-basis code hosphe (v1.02) is, of course, orders of magnitude faster that the 3D code HFODD (v2.40h) [18]. For $N_0 = 36$, the former needs only 20 s of CPU time while the latter needs 250,000 s. Second, for both codes, the dependencies on $N_0$ are clearly given by power lows indicated in the figure. Strangely enough, these power lows are different for $N_0$ below and above $N_0 = 20$. At the moment, no explanation for such a timing pattern could be found.

11. Summary

We have developed a computer code that is able to solve the self-consistent equations resulting from the use of the generalized N3LO energy density functional introduced in Ref. [4]. For different terms of the functional, the resulting mean-field expressions were systematically formulated by using the spherical-tensor representation, which was also introduced in Ref. [4]. In Section 3 we presented the general equations, needed for any choice of basis,
and in Section 4 we developed expressions needed to implement the formalism in the spherical harmonic oscillator basis.

The present version of the program is based on the spherical harmonic oscillator basis and can be used to calculate, for example, ground state properties of spherical nuclei using standard Skyrme forces. For this purpose, it has some advantage compared to other codes, because of its fast execution time and stability.

The motivation to construct the code, and its main new feature, is the possibility to test the influence of higher order derivative terms included in the mean field. The values of the higher order coupling constants can be calculated from density-matrix expansions [19] starting from other interactions or by direct fits to experimental data, and both approaches are presently being explored. The formalism was derived in a general way, so as to include non-spherical multipoles, and thus it can also be used for linear-response calculations [20], which go beyond the mean field. Extensions to include pairing, better treatment of the continuum, and statically deformed densities are also being developed.

Acknowledgements

This work was supported by the Academy of Finland and University of Jyväskylä within the FIDIPRO program and by the Polish Ministry of Science and Higher Education under Contract No. N N 202 328234.

Appendix A. Calculation of coefficients $D_{\tau\rho\kappa\pi}^{U\ell}$ (71)

Coefficients $D_{\tau\rho\kappa\pi}^{U\ell}$ (71) can easily be calculated by using the fact that they exactly correspond to the multiplication of tensors $X$ built from the position vector $x$.

$$X_{\tau\kappa} M_{\rho\pi} = \sum_{U\ell M} CK_{\ell} M_{U} D_{\tau\rho\kappa\pi}^{U\ell} X_{U\ell M}.$$  \hspace{1cm} (A.1)

Tensors $X$ are simply related to spherical harmonics (see Eq. 5.2 (40) in Ref. [5]):

$$X_{\rho\pi M} = \left( \frac{r^{2}}{\sqrt{3}} \right)^{\frac{\rho p}{2}} \frac{4\pi p!}{(2P + 1)!!} Y_{\rho\pi M}(\theta, \phi).$$  \hspace{1cm} (A.2)

Therefore, multiplication law of spherical harmonics (Eq. 5.6(9) in Ref. [5]) gives

$$X_{\tau\kappa} M_{\rho\pi} = \left( -\sqrt{3} \right)^{\frac{\tau u}{2}} r^{u} \left( \frac{4\pi R!}{(2R + 1)!!} \right) \sum_{U\ell M} \frac{(2R + 1)!!}{4\pi (2U + 1)} X_{U\ell M}.$$  \hspace{1cm} (A.3)

and thus the required coefficients read

$$D_{\tau\rho\kappa\pi}^{U\ell} = \delta_{\tau\pi,\rho\kappa} \left( \frac{4\pi R!}{(2R + 1)!!} \right) \frac{U^{1!} U^{1!}(2U + 1)}{(2U - 1)!!(2R - 1)!!} r_{U\rho\pi\ell},$$  \hspace{1cm} (A.4)

and are independent of $urp$.

Appendix B. Derivation of Eq. (97)

We begin by a summation of the product of four Clebsh–Gordan coefficients, that appear in Eq. (96), that is,

$$\sum_{m_{1} m_{2} m_{3} m_{4}} C_{m_{1} m_{2} m_{3} m_{4}}^{\ell_{1} \ell_{2} \ell_{3} \ell_{4}} = \sum_{m_{1} m_{2} m_{3} m_{4}} \left( -1 \right)^{m_{1} + m_{2} + \ell_{1} + \ell_{2}} \left( \frac{2}{2 \ell_{1} + 1} \right)^{2 \ell_{1} + 1} C_{m_{1} m_{2} m_{3} m_{4}}^{\ell_{1} \ell_{2} \ell_{3} \ell_{4}}.$$

$$(\ell_{1} \ell_{2} \ell_{3} \ell_{4}) gives$$

$$= \sum_{m_{1} m_{2} m_{3} m_{4}} \left( -1 \right)^{m_{1} + m_{2} + \ell_{1} + \ell_{2}} \left( \frac{2}{2 \ell_{1} + 1} \right)^{2 \ell_{1} + 1} C_{m_{1} m_{2} m_{3} m_{4}}^{\ell_{1} \ell_{2} \ell_{3} \ell_{4}}.$$

$$= \sum_{m_{1} m_{2} m_{3} m_{4}} \left( -1 \right)^{m_{1} + m_{2} + \ell_{1} + \ell_{2}} \left( \frac{2}{2 \ell_{1} + 1} \right)^{2 \ell_{1} + 1} C_{m_{1} m_{2} m_{3} m_{4}}^{\ell_{1} \ell_{2} \ell_{3} \ell_{4}}.$$

$$= \sum_{m_{1} m_{2} m_{3} m_{4}} \left( -1 \right)^{m_{1} + m_{2} + \ell_{1} + \ell_{2}} \left( \frac{2}{2 \ell_{1} + 1} \right)^{2 \ell_{1} + 1} C_{m_{1} m_{2} m_{3} m_{4}}^{\ell_{1} \ell_{2} \ell_{3} \ell_{4}}.$$

$$= \sum_{m_{1} m_{2} m_{3} m_{4}} \left( -1 \right)^{m_{1} + m_{2} + \ell_{1} + \ell_{2}} \left( \frac{2}{2 \ell_{1} + 1} \right)^{2 \ell_{1} + 1} C_{m_{1} m_{2} m_{3} m_{4}}^{\ell_{1} \ell_{2} \ell_{3} \ell_{4}}.$$

$$= \sum_{m_{1} m_{2} m_{3} m_{4}} \left( -1 \right)^{m_{1} + m_{2} + \ell_{1} + \ell_{2}} \left( \frac{2}{2 \ell_{1} + 1} \right)^{2 \ell_{1} + 1} C_{m_{1} m_{2} m_{3} m_{4}}^{\ell_{1} \ell_{2} \ell_{3} \ell_{4}}.$$

$$= \sum_{m_{1} m_{2} m_{3} m_{4}} \left( -1 \right)^{m_{1} + m_{2} + \ell_{1} + \ell_{2}} \left( \frac{2}{2 \ell_{1} + 1} \right)^{2 \ell_{1} + 1} C_{m_{1} m_{2} m_{3} m_{4}}^{\ell_{1} \ell_{2} \ell_{3} \ell_{4}}.$$

$$= \sum_{m_{1} m_{2} m_{3} m_{4}} \left( -1 \right)^{m_{1} + m_{2} + \ell_{1} + \ell_{2}} \left( \frac{2}{2 \ell_{1} + 1} \right)^{2 \ell_{1} + 1} C_{m_{1} m_{2} m_{3} m_{4}}^{\ell_{1} \ell_{2} \ell_{3} \ell_{4}}.$$

$$= \sum_{m_{1} m_{2} m_{3} m_{4}} \left( -1 \right)^{m_{1} + m_{2} + \ell_{1} + \ell_{2}} \left( \frac{2}{2 \ell_{1} + 1} \right)^{2 \ell_{1} + 1} C_{m_{1} m_{2} m_{3} m_{4}}^{\ell_{1} \ell_{2} \ell_{3} \ell_{4}}.$$

$$= \sum_{m_{1} m_{2} m_{3} m_{4}} \left( -1 \right)^{m_{1} + m_{2} + \ell_{1} + \ell_{2}} \left( \frac{2}{2 \ell_{1} + 1} \right)^{2 \ell_{1} + 1} C_{m_{1} m_{2} m_{3} m_{4}}^{\ell_{1} \ell_{2} \ell_{3} \ell_{4}}.$$

$$= \sum_{m_{1} m_{2} m_{3} m_{4}} \left( -1 \right)^{m_{1} + m_{2} + \ell_{1} + \ell_{2}} \left( \frac{2}{2 \ell_{1} + 1} \right)^{2 \ell_{1} + 1} C_{m_{1} m_{2} m_{3} m_{4}}^{\ell_{1} \ell_{2} \ell_{3} \ell_{4}}.$$

$$= \sum_{m_{1} m_{2} m_{3} m_{4}} \left( -1 \right)^{m_{1} + m_{2} + \ell_{1} + \ell_{2}} \left( \frac{2}{2 \ell_{1} + 1} \right)^{2 \ell_{1} + 1} C_{m_{1} m_{2} m_{3} m_{4}}^{\ell_{1} \ell_{2} \ell_{3} \ell_{4}}.$$

$$= \sum_{m_{1} m_{2} m_{3} m_{4}} \left( -1 \right)^{m_{1} + m_{2} + \ell_{1} + \ell_{2}} \left( \frac{2}{2 \ell_{1} + 1} \right)^{2 \ell_{1} + 1} C_{m_{1} m_{2} m_{3} m_{4}}^{\ell_{1} \ell_{2} \ell_{3} \ell_{4}}.$$
\[ (-1)^{M_1 - M'_1} (-1)^{l + k - u' - T} \times \sqrt{(2k + 1)(2k' + 1)(2T' + 1)(2T + 1)} \times \sum_{W'MWv} \sum_{W'MWv} C_{W'MWv}^{W'MWv} \left\{ \begin{array}{ccc} l & k & U \\ l' & k' & U' \\ T & T' & W \end{array} \right\}, \] 

(B.3)

see Eq. 8.7(20) in Ref. [5].

We can now perform the summation over \( M'_1 \) and \( M_1 \), which gives the factor \((-1)^{U + U' - W} b_{\beta\gamma}^{\delta}\delta_{M_1'\gamma}M_1\delta_{M_W,M_W'}\) and allows for a summation over \( W' \) and \( M_W' \). After inserting all these results into Eq. (96), we obtain

\[ \tilde{\rho}_{v'jf'M'} = \sum_{MWv'MWv'M} \sum_{W'MWv'M} C_{W'MWv'M}^{W'MWv'M} \left\{ \begin{array}{ccc} l & k & U \\ l' & k' & U' \\ T & T' & W \end{array} \right\}. \] 

(B.4)

Now we have to sum up products of three Clebsh–Gordan coefficients:

\[ \sum_{MW_1'M_1W_2'M_2W_3'M_3} (-1)^{M_1'} C_{W_1'M_1W_2'M_2W_3'M_3}^{W_1'M_1W_2'M_2W_3'M_3} \left\{ \begin{array}{ccc} l & k & U \\ l' & k' & U' \\ T & T' & W \end{array} \right\}. \] 

(B.5)

see Eq. 8.7(15) in Ref. [5]. This gives:

\[ \tilde{\rho}_{v'jf'M'}^{u'W'f'} = \] 

\[ (-1)^{U} \sum_{Nlj} b^{3 + 3 + u' + u''} e^{(-u'' - u' - T)} \times \sum_{W'TM} \sum_{W'TM} C_{W'TM}^{W'TM} \left\{ \begin{array}{ccc} l & k & U \\ l' & k' & U' \\ T & T' & W \end{array} \right\} \] 

(B.6)

Finally we have:

\[ \tilde{\rho}_{v'jf'M'}^{u'W'f'} = \] 

\[ (-1)^{1 + v + l + k - u' - W} \] 

where we have also used Eq. (89). This gives expression (97) and definitions of radial form factors (98) and coefficients (99).

Appendix C. Derivation of Eq. (112)

After inserting Eq. (110) into Eq. (111) we must sum up products of four Clebsh–Gordan coefficients, similarly as in Eq. (81).
that is,
\[
\sum_{m_j m_{\ell} m_{I}} c_{\ell m_{\ell}}^{m_j} c_{m_{\ell} m_{I} m_{I}}^{m_j} c_{m_{I} m_{I} m_{I}}^{m_j} c_{m_{I} m_{I} m_{I}}^{m_j} = (-1)^{M'_{I} - m_{-I'_{I}} - v'} \sqrt{2/2j' + 1} \sqrt{2/2j' + 1} \sqrt{2/2j' + 1} \\
\times \sum_{T'} T' T_m' T_{m'_{I}} \left\{ j \quad j' \quad j'' \right\} \left\{ l \quad i \quad l' \quad i' \quad i'' \right\} \left\{ T' \quad T \quad W' \right\},
\]
(C.1)
and then as in Eq. (B.3):
\[
\sum_{m_j m_{\ell} m_{I}} (-1)^{m_j} c_{-m_j m_{\ell} m_{I}} c_{m_{I} m_{I} m_{I}}^{m_j} c_{m_{I} m_{I} m_{I}}^{m_j} c_{m_{I} m_{I} m_{I}}^{m_j} = (-1)^{l + k' + k'} \sqrt{(2k + 1)(2k' + 1)(2k' + 1)(2T' + 1)} \\
\times \sum_{W' M_{W}} (-1)^{M'_{W} - M_{W} m_{W}} c_{W' M_{W}}^{M'_{W}} c_{W' M_{W}}^{W' M_{W}} c_{W' M_{W}}^{W' M_{W}} c_{W' M_{W}}^{W' M_{W}} \left\{ l \quad k \quad l' \quad k' \quad l'' \right\} \left\{ T' \quad T \quad W' \right\}.
\]
(C.2)

We can now perform the summation over $M'_{I}$ and $M_{I}$, which gives the factor $(-1)^{l + f - L'} \delta_{M'_{I} M_{I}}$, and allows for a summation over $W'$ and $M_{W}$. After inserting all these results into Eq. (111), we obtain
\[
\langle \Phi_{N'T'} | \hat{H}^{\mu} | \langle \beta J^M' \rangle | \Phi_{NI} \rangle
\]
\[
= \sum_{M'_{I} M_{I}} \sum_{W' M_{W}} \sum_{m_{-I'_{I}} m_{I} m_{I}} \sum \int r^2 \ dr \\
\times C_{M'_{I} M_{I} M_{I} M_{I}}^{I M_{I} M_{I} M_{I}} (-1)^{M'_{I} - M_{I} + 1} (-1)^{m_{I} m_{I}} C_{W_{M_{W}}}^{W' M_{W}} C_{W_{M_{W}}}^{W' M_{W}} C_{W_{M_{W}}}^{W' M_{W}} C_{W_{M_{W}}}^{W' M_{W}} \left\{ l \quad k \quad l' \quad k' \quad l'' \right\} \left\{ T' \quad T \quad W' \right\}.
\]
(C.3)

Now, we have to sum up products of three Clebsch–Gordan coefficients, similarly as in Eq. (B.5), that is,
\[
\sum_{M'_{I} M_{I} M_{I} M_{I}} (-1)^{M'_{I} M_{I} M_{I} M_{I}} c_{M'_{I} M_{I} M_{I} M_{I}}^{I M_{I} M_{I} M_{I}} c_{M'_{I} M_{I} M_{I} M_{I}}^{I M_{I} M_{I} M_{I}} c_{M'_{I} M_{I} M_{I} M_{I}}^{I M_{I} M_{I} M_{I}} c_{M'_{I} M_{I} M_{I} M_{I}}^{I M_{I} M_{I} M_{I}} = \sqrt{2T' + 1} \sqrt{2T' + 1} \sqrt{2T' + 1} \sqrt{2T' + 1} \\
\times (-1)^{l + f - L'} \sqrt{2T' + 1} \sqrt{2T' + 1} \sqrt{2T' + 1} \sqrt{2T' + 1} \\
\times \sum_{T' M_{I}} c_{T' M_{I} M_{I} M_{I}}^{T' M_{I} M_{I} M_{I}} \left\{ l \quad k \quad l' \quad k' \quad l'' \right\} \left\{ T' \quad T \quad L' \right\}.
\]
(C.4)

This gives:
\[
\langle \Phi_{N'T'} | \hat{H}^{\mu} | \langle \beta J^M' \rangle | \Phi_{NI} \rangle
\]
\[
= \sum_{M'_{I} M_{I}} \sum_{W' M_{W}} \sum_{m_{-I'_{I}} m_{I} m_{I}} \sum \int r^2 \ dr \sum_{k} \frac{1}{\sigma_{\nu}} \left\{ l \quad k \quad l' \quad k' \quad l'' \right\} \left\{ T' \quad T \quad W' \right\}.
\]
(C.5)

Finally, summations over $M'_{I}$ and $M_{I}$ give the factor $\delta_{l + f - L'} \delta_{M'_{I} M_{I}}$, that is:
\[
\langle \Phi_{N'T'} | \hat{H}^{\mu} | \langle \beta J^M' \rangle | \Phi_{NI} \rangle
\]
\[
= \delta_{l + f - L'} \sum_{W' M_{W}} \sum_{m_{-I'_{I}} m_{I} m_{I}} \sum \int r^2 \ dr b_{W_{M_{W}}}^{\nu} F_{m_{-I'_{I}} m_{I} m_{I} M_{I}}^{\nu} (br) F_{m_{-I'_{I}} m_{I} m_{I} M_{I}}^{\nu} (br) e^{-2(2r)^2} \\
\times C_{T' M_{I} M_{I} M_{I}}^{T' M_{I} M_{I} M_{I}} \left\{ l \quad k \quad l' \quad k' \quad l'' \right\} \left\{ T' \quad T \quad W' \right\}.
\]
(C.6)

where we have also used Eq. (89). This shows explicitly that in the field calculated for the density matrix with multipolarity $J''$, only the multipole $J''$ appears.

References

